Study Dark Matter & Cosmic Neutrinos while Defending the Earth

Yu-Dai Tsai

University of California, Irvine With Josh Eby, Jason Arakawa, Marianna Safronova Youjia Wu, Sunny Vagnozzi, Luca Visinelli Contact: yudait1@uci.edu & yt444@cornell.edu Cornell -> Fermilab -> UC Irvine

Parker Solar Probe

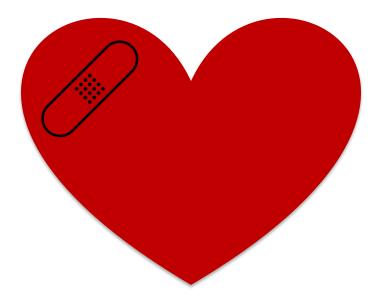
Credit: NASA/Johns Hopkins APL/Steve Gribben

<u>https://arxiv.org/abs/2112.07674</u>
 <u>https://arxiv.org/abs/2107.04038</u>
 <u>Under review by Nature Astronomy</u>

My outreach interview: <u>https://www.youtube.com/watch?v=xDX9XwLHBuM</u> (> 76K views!) I will first have the mental-health discussion, as it is very important!

Hidden Sector of Diversity

- 1. Normalize the discussion of mental health
- 2. Understand & support invisible disabilities
- 3. Value neurodiversity



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Academia's Mental Health Crisis

- We are doing way worse than national average
- We cannot wait for a national outcry to improve this issue

"41% of graduate students scored as having moderate to severe anxiety on the GAD07 scale as compared to 6% of the general population", see https://www.nature.com/articles/nbt.4089, Evans et al.

- Many other studies support this.
- It is not hard to improve it: implement useful changes

My Definitions for Discussions

- Invisible disabilities: diagonalizable disabilities including ADHD, PTSD, autism, learning disabilities, chronic pain (often pre-existing; not visible)
- Mental health issues: general behavioral or mental pattern that causes distress or impairment of personal functioning
- **Neurodiversity**: variations in the human brain and cognition, for instance in sociability, learning, attention, mood, and other mental functions

My journey

- I have 55 publications including research & white papers;
 roughly 1,400 citations. But being different, I often feel like an outsider
- Taiwan had serious child abuse issues in school, conducted by teachers & normalized in society
- I overcame untreated invisible disabilities; but without treating them & proper support, my life expectancy, personal functioning, and relationships were greatly impaired
- It's not enough to treat them. We also need understandings.
- I realized that I can be a unique voice in our community

Open-up is difficult, but very helpful!

- I benefit from a senior/respected collaborator open-up about the personal struggles in social media
- It finally urged me to face my own invisible disabilities & I received medical treatments

• Life is so much better. I now have longer life expectancy & satisfaction (backed by studies!)

Aspen Mental Health Lunch

The lunch is confidential, but one can cover topics include

- 1. Utilize resource inc. HR, lab/school sources, etc.
- 2. Develop mentoring & group meeting plans
- 3. Increase representation of neurodiverse people; Reduce negative connotation/descriptions
- 4. Conflict resolution for neuro-diverse people
- 5. Co-learning with LGBTQIA community
- 6. Intersectionality: gender, racial, cultural discussionsI am happy to be your guest host such a lunch in yourdepartment :)

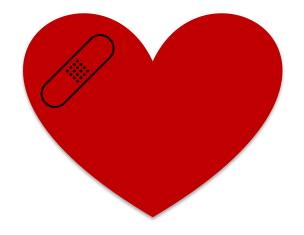
Important Efforts

1. A. Welsh, "Mental health for academics;" Slack channel

 Assamagan, Bitter, Chen, Choi, Esquivel, Jepsen, Lewis, Muronga, Walkowicz, Zhang, Snowmass Whitepaper: "Building a Culture of Equitable Access and Success for Marginalized Members in Today's Particle Physics Community," <u>https://arxiv.org/abs/2206.01849</u> Specifically discussed invisible disability!

3. S. Charley, "<u>How to be human in physics</u>," looking for your stories to share.

Let us all



- 1. Normalize the discussion of mental health
- 2. Support & understand invisible disabilities
- 3. Value neurodiversity

I encourage your own neuro-diverse ways to improve our community!

- My vision is one day we can all open up about our struggles without the fear of discrimination
- Follow mental-health science
- Increase representation of people with invisible disabilities
- Reduce nepotism & <u>support independence</u>: an academia for everyone

Yu-Dai Tsai, UC Irvine, yudait1@uci.edu & yt444@cornell.edu

We can also discus & collaborate on ...

- Fixed-target searches for dark matter & long-lived particles (FerMINI & LongQuest) with Pospelov et al.
- LHC Forward Experiments: Forward Physics Facility, **FORMOSA** (a millicharge experiment I proposed), with Feng et al.
- Dark matter model building (dark sector QCD, Strongly Self-Interacting Dark Matter, SIMP/ELDER), with Murayama, Slatyer, Perelstein et al.
- Dark matter searches using neutron star / compact merger / multimessenger astronomy, with Profumo, Sathyaprakash et al.
- Neutrino physics (cosmic neutrino background) & neutrino BSM, with Shoemaker et al.
- Collaborating with **many awesome early-career collaborators.**

Invisible disabilities cultivate diverse abilities

Planetary Defense & Space Quantum Technologies for Fundamental Physics

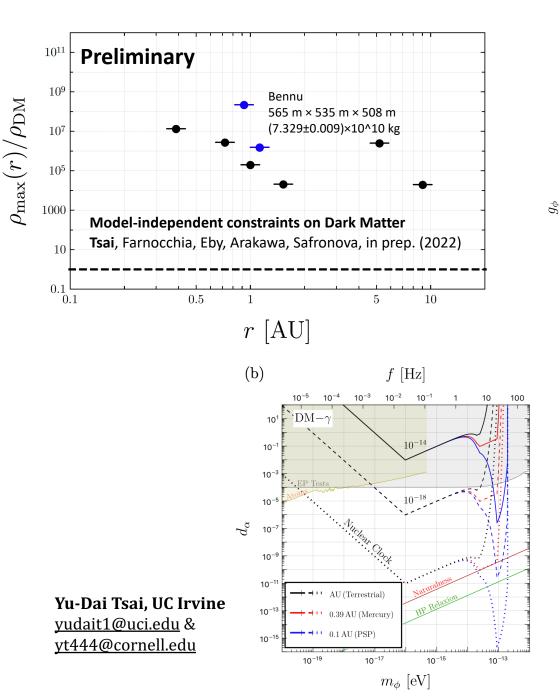
Yu-Dai Tsai University of California, Irvine With Josh Eby, Jason Arakawa, Marianna Safronova Youjia Wu, Sunny Vagnozzi, Luca Visinelli Contact: yudait1@uci.edu & yt444@cornell.edu Cornell -> Fermilab -> UC Irvine

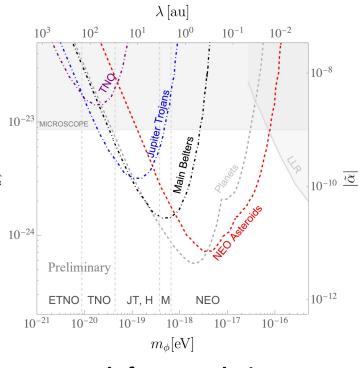
Parker Solar Probe

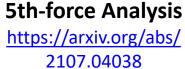
Credit: MASA/Johns Hopkins APL/Steve Gribben

https://arxiv.org/abs/2112.07674
 https://arxiv.org/abs/2107.04038
 Submitted to Nature Astronomy

My outreach interview: <u>https://www.youtube.com/watch?v=xDX9XwLHBuM</u> (> 75K views!)







Ultralight DM & Quantum Sensor

https://arxiv.org/abs/2112.07674

Big Questions

Can planetary data set meaningful constraints on
 Dark matter?
 General Relativity?
 5th forces?



 Can we use current or future space quantum technologies to study fundamental physics?

Yu-Dai Tsai, UC Irvine, yudait1@uci.edu & yt444@cornell.edu

Answers

- Can planetary data set meaningful dark matter constraints?
 General Relativity?
 5th forces?
 Yes! Many opportunities
- Can we use current or future space quantum technology to study fundamental physics?
 Yes, I will show you an example today.
- Robust analyses underway in collaboration with a NASA JPL planetary defense expert & an AMO expert

Outline

- Planetary Defense & Fifth Forces Studies
- Model-Independent Probes of ANY Dark Matter Candidates (especially purely gravitational dark matter!)
- Space Quantum Clocks & Ultralight Dark Matter
- New Projects & Visions for the Future

Theme of this talk:

Bridging Planetary Science, Space/Quantum Technologies, and Fundamental Physics

Many real-life applications & consequences!



Self-driving Spacecraft!

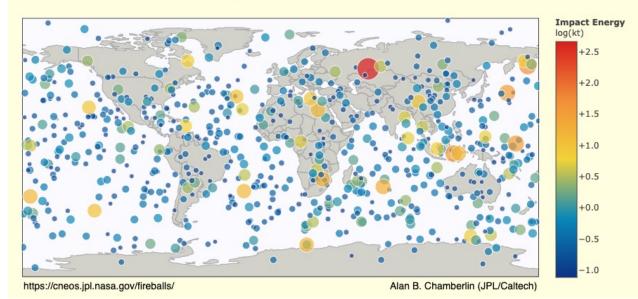
Please Look Up

Planetary Constraints: Dark Matter & Fifth Forces

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Asteroids hitting the earth

Fireballs Reported by US Government Sensors (1988-Apr-15 to 2021-Jul-30)



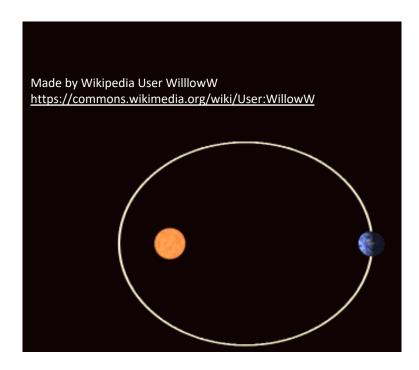


 ~ 65 million years ago

Tracking asteroids is extremely important e.g., unexpected 2013 Chelyabinsk meteor injured >1500 people Also, near-Earth asteroid search accidentally found 'Oumuamua

Perihelion Precession: Einstein's Success

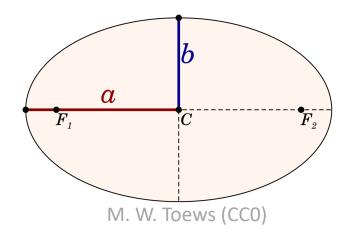
Precession of Mercury's perihelion (closest point to the Sun)



https://en.wikipedia.org/wiki/Apsidal_precession#/media/File:Prec essing_Kepler_orbit_280frames_e0.6_smaller.gif under CC BY 3.0

$$\frac{\mathrm{d}^2 u}{\mathrm{d}\varphi^2} + u - \frac{GM_{\odot}}{L^2} = \frac{3GM_{\odot}}{c^2} u^2 \cdot \mathbf{GR}$$

- Consider planar motion and fix $\theta = \pi/2$.
- Define inverse radius variable $u \equiv 1/r = u(\phi)$
- $a = \frac{L^2}{M_{\odot}(1-e^2)}$, a is the semi-major axis



Objects of interest

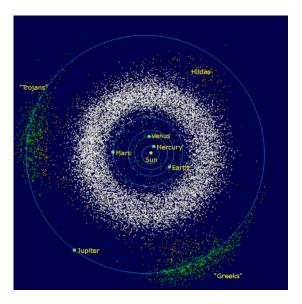
Minor Planets	a [au]	\sim Numbers
Near-Earth Object (NEO)	$< 1.3^*$	> 25000
Main-Belt Asteroid (M)	$\sim 2-3$	~ 1 million
Hilda (H)	3.7 - 4.2	> 4000
Jupiter Trojan (JT)	5.2	> 9800
Trans-Neptunian Object (TNO)	> 30	2700
Extreme TNO (ETNO)	> 150	12

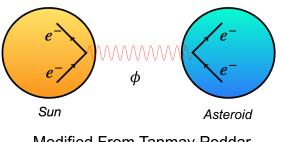
TABLE I. Targets for our future studies, for which exciting opportunities are provided by sheer numbers and observational programs, classified roughly based on their typical semimajor axes.

*NEOs are defined as having perihelia a(1 - e) < 1.3 au.

$$|\Delta \varphi_{\phi,A'}| \simeq rac{2\pi}{1+rac{g^2}{4\pi G m_p^2}} rac{g^2}{4\pi G m_p^2} \left(rac{amc}{\hbar}
ight)^2 (1-\mathsf{e}) \,.$$

- Tsai, Wu, Vagnozzi, Visinelli, <u>arXiv:2107.04038</u>
- Can also probe dark matter, primordial black hole, etc





Modified From Tanmay Poddar

Ultralight Mediators & Fifth Forces

Gauged $U(1)_{EM}$ (Standard Model) \Rightarrow photons

1. Gauged $U(1)_{X's}$ (hypothetical) \implies "Dark" photons

- X can be bayon number, lepton number, etc: Standard Model Global Symmetries
- Motivated by baryogenesis (matter-anti matter asymmetry) & dark matter:

The ultralight mediators **CAN** but does not have to be dark matter

2. Ultralight scalars coupled to Standard Model particles

 $\mathcal{L}_{\phi} \subset (g_{\phi,p} ar{p} p \ + \ g_{\phi,n} ar{n} n \ + \ g_{\phi,e} ar{e} e \) \phi$

5th force and Yukawa Potential

$$\begin{split} V(r) &= \widetilde{\alpha} \frac{GM_{\odot}M_{*}}{r} \exp\left(-\frac{r}{\lambda}\right) \,, \\ V(r) &= \mp \frac{g^{2}}{4\pi} \frac{Q_{\odot}Q_{*}}{r} \exp\left(-\frac{mc^{2}}{\hbar c}r\right) \,, \\ \frac{d^{2}u}{d\varphi^{2}} &+ u - \frac{GM_{\odot}}{L^{2}} = \frac{3GM_{\odot}}{c^{2}}u^{2} + \widetilde{\alpha} \frac{GM_{\odot}}{L^{2}} \left(1 + \frac{1}{\lambda u}\right)e^{-\frac{1}{\lambda u}} \,, \end{split}$$
(fifth force)

- Gauge boson, dark photon of $U(1)_B$ or scalar coupled to baryon number
- g is new physics coupling constant, and m is the mediator mass
- See, e.g., Poddar et al, https://arxiv.org/abs/2002.02935

Precession (Analytical) at Low-Mass Limit

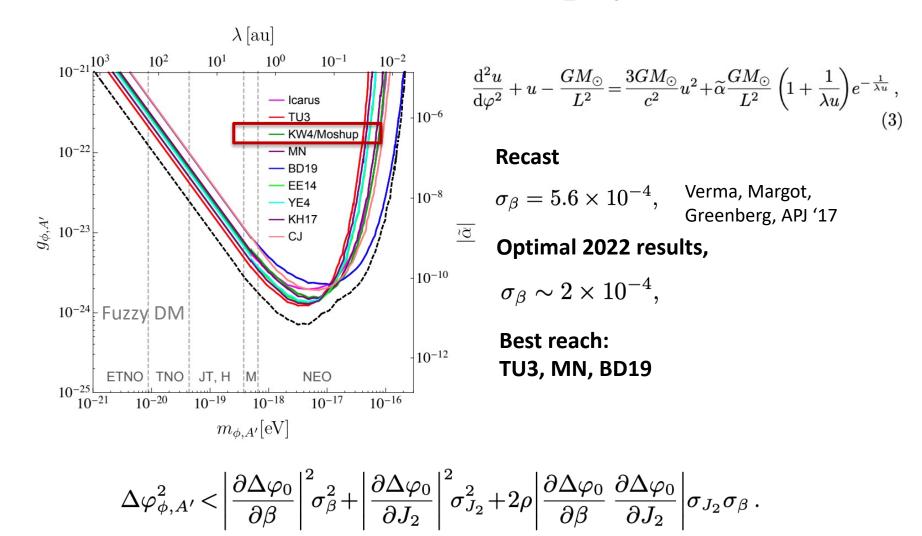
$$\begin{split} |\Delta\varphi_{\phi,A'}| \simeq \frac{2\pi}{1 + \frac{g^2}{4\pi G m_p^2}} \frac{g^2}{4\pi G m_p^2} \left(\frac{amc}{\hbar}\right)^2 (1 - \mathbf{e}) \,. \end{split}$$
 (fifth force)

• m_p is proton mass

$$\Delta \varphi_0 = \frac{6\pi G M_{\odot}}{a(1-\mathsf{e}^2)c^2} \left[\frac{2-\beta+2\gamma}{3}\right]$$
(GR)

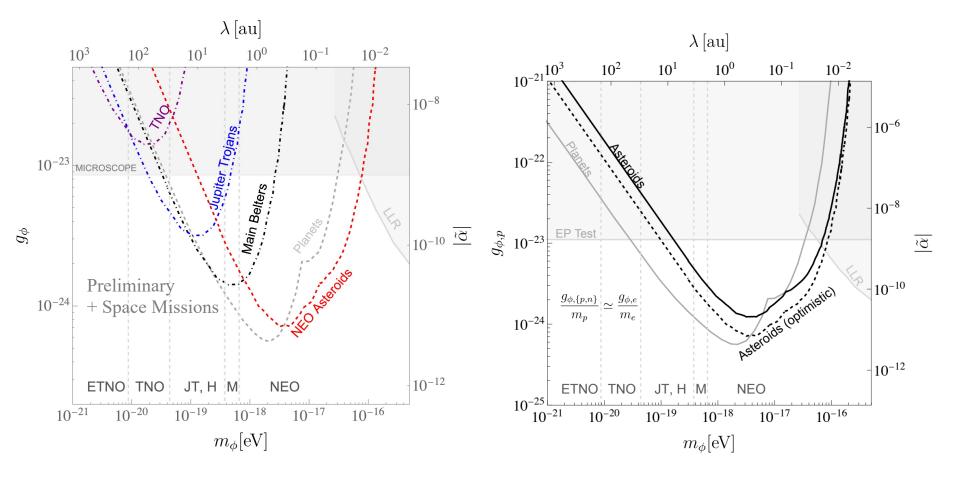
- for low mass, m << 1/ a (Natural Unit)
- The term gets larger with *a*
- That's why we should explore **objects further away from the Sun:** not just Mercury or other planets
- Not depending on target celestial bodies' mass

Results for the new physics



Tsai, Wu, Vagnozzi, Visinelli, <u>arXiv:2107.04038</u>

Compilations of Various Probes

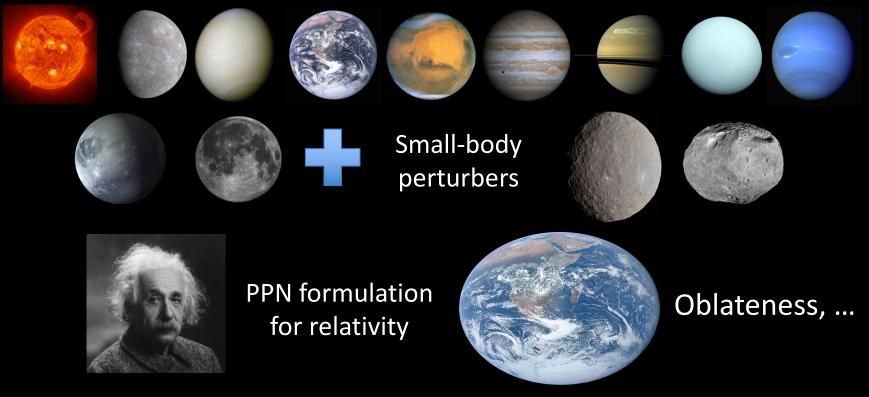


Tsai, Wu, Vagnozzi, Visinelli, <u>arXiv:2107.04038</u>

We are conducting a **detailed study** using **MONTE** with people from JPL & ESA 25

Robust Analysis: High-fidelity force model

JPL Planetary Ephemerides DE441



Dr. Davide Farnocchia's (NASA, JPL) slide

Adding Fifth Forces or Dark Matter to the Force Model

$$\begin{split} \ddot{\mathbf{r}}_{i} &= \sum_{j \neq i} \frac{\mu_{j} \left(\mathbf{r}_{j} - \mathbf{r}_{i} \right)}{r_{ij}^{3}} \begin{cases} 1 - \frac{2(\beta + \gamma)}{c^{2}} \sum_{l \neq i} \frac{\mu_{l}}{r_{il}} - \frac{2\beta - 1}{c^{2}} \sum_{k \neq j} \frac{\mu_{k}}{r_{jk}} \\ &+ \gamma \left(\frac{\dot{s}_{i}}{c} \right)^{2} + (1 + \gamma) \left(\frac{\dot{s}_{j}}{c} \right)^{2} - \frac{2(1 + \gamma)}{c^{2}} \dot{\mathbf{r}}_{i} \cdot \dot{\mathbf{r}}_{j} \\ &- \frac{3}{2c^{2}} \left[\frac{\left(\mathbf{r}_{i} - \mathbf{r}_{j} \right) \cdot \dot{\mathbf{r}}_{j}}{r_{ij}} \right]^{2} + \frac{1}{2c^{2}} \left(\mathbf{r}_{j} - \mathbf{r}_{i} \right) \cdot \ddot{\mathbf{r}}_{j} \\ &+ \frac{1}{c^{2}} \sum_{j \neq i} \frac{\mu_{j}}{r_{ij}^{3}} \left\{ \left[\mathbf{r}_{i} - \mathbf{r}_{j} \right] \cdot \left[(2 + 2\gamma) \dot{\mathbf{r}}_{i} - (1 + 2\gamma) \dot{\mathbf{r}}_{j} \right] \right\} \left(\dot{\mathbf{r}}_{i} - \dot{\mathbf{r}}_{j} \right) \\ &+ \frac{3 + 4\gamma}{2c^{2}} \sum_{j \neq i} \frac{\mu_{j} \ddot{\mathbf{r}}_{j}}{r_{ij}} \end{split}$$

From Dr. Davide Farnocchia's (NASA, JPL) slide

F =
$$\frac{A0 e^{-\frac{r}{r_0}}}{r^2} + \frac{A0 e^{-\frac{r}{r_0}}}{r r^0}$$

General dark matter

Uncertainties?

- Errors in planetary trajectories and masses
- Missing perturbers, errors in perturber masses & trajectories
- Higher order relativistic terms
- Higher order gravity terms
- Poynting-Robertson drag
- Simplifying assumptions in nongravitational force model (non-spherical effects, Yarkovsky,

solar torque, physical parameter evolution, etc)

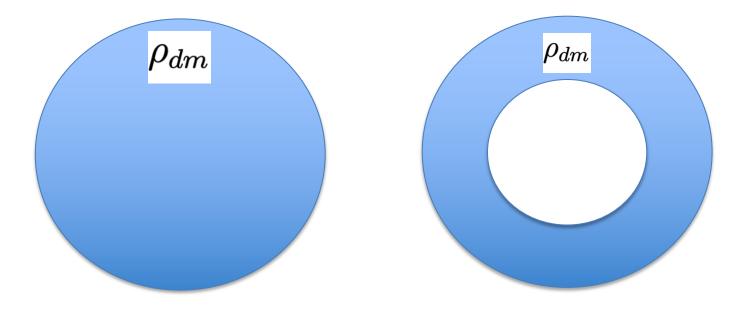
- Solar mass loss and solar wind
- Meteoroid impacts
- Spacecraft interaction
- Whatever else could be missing...

Dr. Davide Farnocchia's (NASA, JPL) slide

Model-Independent Constraints on Dark Matter Preliminary Results

Yu-Dai Tsai, UC Irvine, yudait1@uci.edu & yt444@cornell.edu

Dark Matter Profile & Planetary Precession



 $\Delta\theta_0 = -4\pi^2 \rho_{dm} a^3 (1-e^2)^{1/2} / M_{\odot},$

Adding Dark Matter to the Force Model

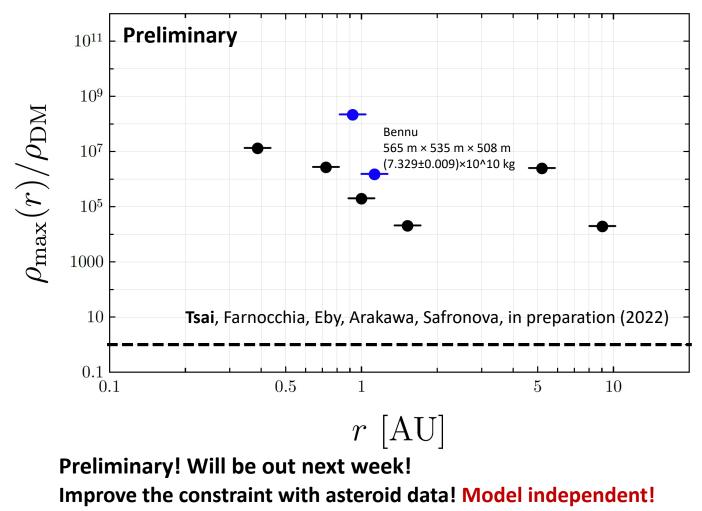
$$\begin{split} \ddot{\mathbf{r}}_{i} &= \sum_{j \neq i} \frac{\mu_{j} \left(\mathbf{r}_{j} - \mathbf{r}_{i} \right)}{r_{ij}^{3}} \left\{ 1 - \frac{2(\beta + \gamma)}{c^{2}} \sum_{i \neq i} \frac{\mu_{l}}{r_{il}} - \frac{2\beta - 1}{c^{2}} \sum_{k \neq j} \frac{\mu_{k}}{r_{jk}} \right. \\ &+ \gamma \left(\frac{\dot{s}_{i}}{c} \right)^{2} + (1 + \gamma) \left(\frac{\dot{s}_{j}}{c} \right)^{2} - \frac{2(1 + \gamma)}{c^{2}} \dot{\mathbf{r}}_{i} \cdot \dot{\mathbf{r}}_{j} \\ &- \frac{3}{2c^{2}} \left[\frac{\left(\mathbf{r}_{i} - \mathbf{r}_{j} \right) \cdot \dot{\mathbf{r}}_{j}}{r_{ij}} \right]^{2} + \frac{1}{2c^{2}} \left(\mathbf{r}_{j} - \mathbf{r}_{i} \right) \cdot \ddot{\mathbf{r}}_{j} \right] \\ &+ \frac{1}{c^{2}} \sum_{j \neq i} \frac{\mu_{j}}{r_{ij}^{3}} \left\{ \left[\mathbf{r}_{i} - \mathbf{r}_{j} \right] \cdot \left[(2 + 2\gamma) \dot{\mathbf{r}}_{i} - (1 + 2\gamma) \dot{\mathbf{r}}_{j} \right] \right\} \left(\dot{\mathbf{r}}_{i} - \dot{\mathbf{r}}_{j} \right) \\ &+ \frac{3 + 4\gamma}{2c^{2}} \sum_{j \neq i} \frac{\mu_{j} \ddot{\mathbf{r}}_{j}}{r_{ij}} \end{split}$$

From Dr. Davide Farnocchia's (NASA, JPL) slide



$$\begin{split} F(r) &= \frac{2\pi}{3} Gm\rho_0 \left(\frac{2r_0^3}{r^2} - 2r\right) \mathbf{\hat{r}} \\ &\simeq -\frac{4\pi}{3} Gm\rho_0 r\mathbf{\hat{r}} \end{split}$$

New Project: New Model Independent Constraints!



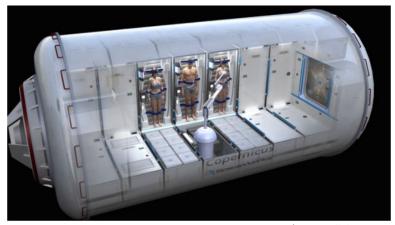
Tsai, Farnocchia, Eby, Arakawa, Safronova, in preparation **Obtained preliminary results from NASA JPL code.**

More on Space Missions & Quantum Sensors

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Why Space Quantum Clocks? Auto-Navigating Spacecraft & Space Travel





Artist's concept for Mars-ready habitat. Image Credit: SpaceWorks Torpor/NASA collaboration

Exploring the deep space: auto-driving Spacecraft; needs precision timing!!!

NASA Deep Space Atomic Clocks (current technology!) & Deep space and global navigation satellite system (GNSS)

Can we use the technology to study fundamental physics?

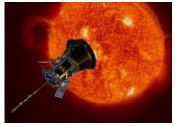
NASA DSAC & Parker Solar Probe



- Deep Space Atomic Clock loses one second every 10 million years, as proven in controlled tests on Earth.
- The clock has operated for more than 12 months in space; demonstrated long-term fractional frequency stability of 3 × 10⁻¹⁵

Burt, Prestage, Tjoelker, Enzer, Kuang, Murphy et al., Nature 595 (2021) 43.

• Exceeds previous space clock performance by up to an order of magnitude



 $(1.0 \text{ m} \times 3.0 \text{ m} \times 2.3 \text{ m})$

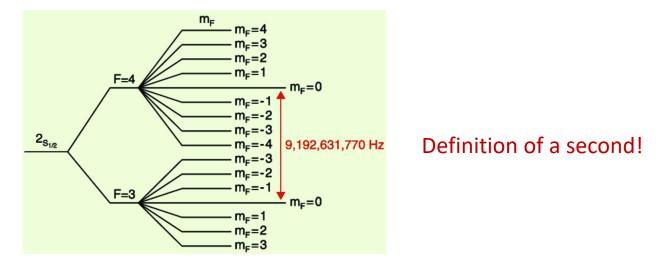
Parker Solar Probe

Kasper, Klein, Lichko, Huang, Chen, Badman et al., Parker solar probe enters the magnetically dominated solar corona, Phys. Rev. Lett. (2021)

Why don't we put a quantum clock on a solar probe?
 What can we do with that?

Atomic Clock & Caesium Standard

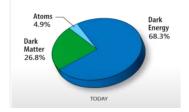
- Atomic clocks: used to measure the distance between objects by timing how long it takes a signal to travel from A to B.
- For space exploration, clocks must be extremely precise:
- An error of even one second can mean the difference between landing on Mars or missing it by hundreds of thousands of miles.



http://hyperphysics.phy-astr.gsu.edu/hbase/acloc.html Reference: U.S. Naval Observatory, Cesium Clocks

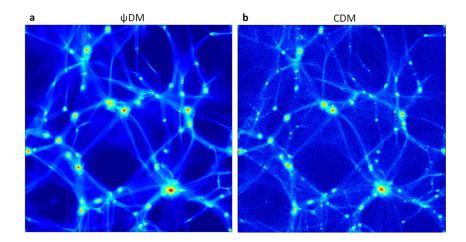
Wave-Like Particles as Dark Matter

$$\lambda_{\rm dB} \equiv \frac{2\pi}{mv}$$
$$N_{\rm dB} \sim \left(\frac{34\,{\rm eV}}{m}\right)^4 \left(\frac{250\,{\rm km/s}}{v}\right)^3 \text{ in } \lambda_{\rm dB}^3$$



UC Riverside Physics Department https://physics.ucr.edu/image/dar k-matter-dark-energy-pie-chart

- For m << 30 eV, the occupancy NdB is so large that the particles are best described by classical waves
- like electromagnetism, a state with a large number of photons is described by the classical EM fields.



Schive, Chiueh, Broadhurst, Nature Physics '14 arXiv:1406.6586,

Oscillation of Wave-like Scalars

$$V(\phi)=rac{1}{2}m_{\phi}^2\phi^2+rac{1}{3}a_{\phi}\phi^3+rac{1}{4}\lambda_{\phi}\phi^4.$$
Dark matter potential

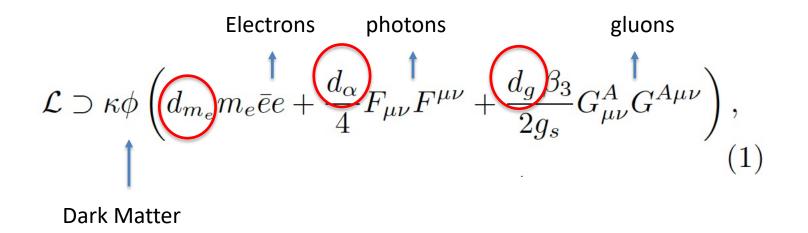
$$\phi(t, \vec{x}) = \phi_0 \cos(m_\phi t - \vec{k}_\phi \cdot \vec{x} + \dots).$$

(Non-relativistic solutions)

 $\omega \simeq m_{\phi}.$

Oscillation frequency ~ dark matter mass

Dark Matter Coupling



where e is the electron field, $F^{\mu\nu}$ ($G^{A\mu\nu}$) is the electromagnetic (QCD) field strength, $|g_s|$ and β_3 are the strong interaction coupling constant and beta function (respectively), and $\kappa = \sqrt{4\pi}/M_P$ with $M_P = 1.2 \times 10^{19}$ GeV.

Atomic Physics Probe

$$\mathcal{L} \supset \kappa \phi \left(d_{m_e} m_e \bar{e} e + \frac{d_\alpha}{4} F_{\mu\nu} F^{\mu\nu} + \frac{d_g \beta_3}{2g_s} G^A_{\mu\nu} G^{A\mu\nu} \right),$$
(1)

$$\mu(\phi) \simeq \mu_0 \left(1 + d_{m_e} \kappa \phi \right), \quad \alpha(\phi) \simeq \alpha_0 \left(1 - d_\alpha \kappa \phi \right)$$
$$\alpha_s(\phi) \simeq \alpha_{s,0} \left(1 - \frac{2d_g \beta_3}{g_s} \kappa \phi \right), \quad (2)$$

where $\mu = m_e/m_p$ is the electron-proton mass ratio, and the subscript $_0$ denotes the central (time-independent) value of μ , α , and α_s .

Atomic Probe Basics

$$\mathcal{L} \supset \kappa \phi \left(d_{m_e} m_e \bar{e} e + \frac{d_\alpha}{4} F_{\mu\nu} F^{\mu\nu} + \frac{d_g \beta_3}{2g_s} G^A_{\mu\nu} G^{A\mu\nu} \right),$$
(1)

Turning off d_{m_e} and d_g for demonstrations,

 $f_A \propto \alpha^{\xi_A+2}$, f is the frequency of a (clock) transition.

$$\alpha = \alpha_0 (1 + d_\alpha \kappa \phi(t)).$$

$$\frac{\delta(f_A/f_B)}{f_A/f_B} \simeq (\xi_A - \xi_B) d_\alpha \kappa \phi(t).$$

- Experimental observable! See arXiv:1405.2925, Arvanitaki, Huang, Tilburg, PRD 15
- For example, if A is a hyperfine microwave transition and B is an electronic optical transition, ζA = 1 and ζB = 0.
- Clock (~ 10⁻¹⁵ for DSAC) stability translate to how well we can measure $\frac{\delta(f_A/f_B)}{f_A/f_B}$

Solar Bound-State Halo

Yu-Dai Tsai, UC Irvine, yudait1@uci.edu & yt444@cornell.edu

Scalar DM Halo

Stable solution can be supported by external potential

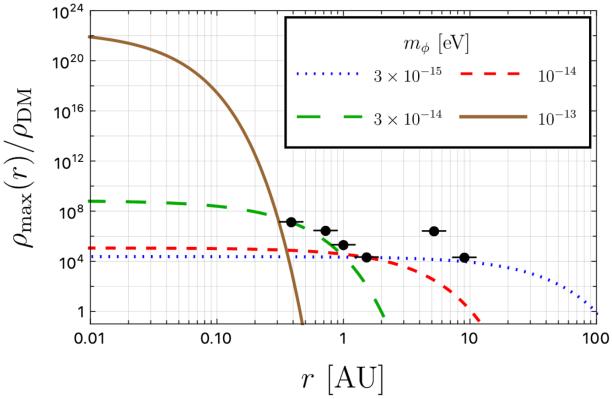
$$V_{\text{ext}} = \begin{cases} -\frac{G \, m_{\phi} \, M_{\text{ext}}}{r} & \text{for } R_{\star} > R_{\text{ext}} ,\\ -\frac{3 \, G \, m_{\phi} \, M_{\text{ext}}}{2 \, R_{\text{ext}}} \left[1 - \frac{1}{3} \left(\frac{r}{R_{\text{ext}}} \right)^2 \right] & \text{for } R_{\star} \le R_{\text{ext}} , \end{cases}$$

$$ho(r)\simeq
ho_\star\exp\left(-2r/R_\star
ight)$$
, for $R_\star>R_{
m ext}$

 $R_{\star} \simeq \frac{M_P^2}{M_{\text{ext}} m_{\phi}^2},$ where $M_{\text{ext}} = M_{\odot}$ is the mass of the external host body; note that R_{\star} is independent of the total mass in the halo $v_{\star} = (m_{\phi} R_{\star})^{-1},$

Banerjee, Budker, Eby, Flambaum, Kim, Matsedonskyi, and Perez, 1912.04295

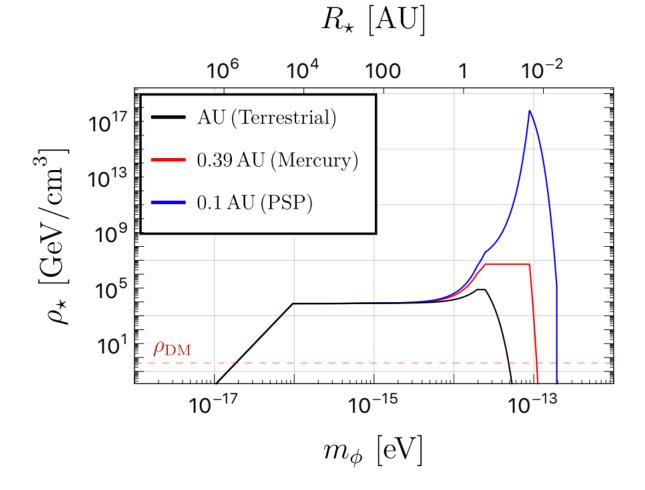
Dark matter in solar system? Planetary constraint!



• Black data points are model-independent constraints!

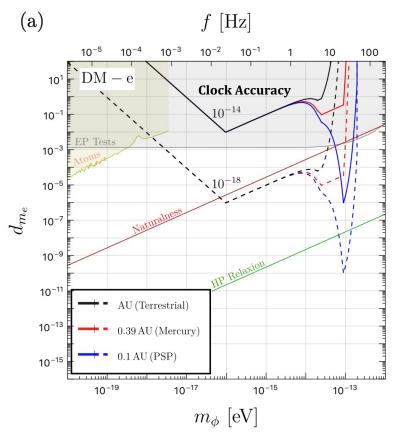
• Dark matter induce precessions to the planets Mercury, Venus, Earth, Mars, Jupiter, Saturn Pitjev, Pitjeva, 1306.5534, Astronomy Letters '13 Tsai, Eby, Safronova, 2112.07674

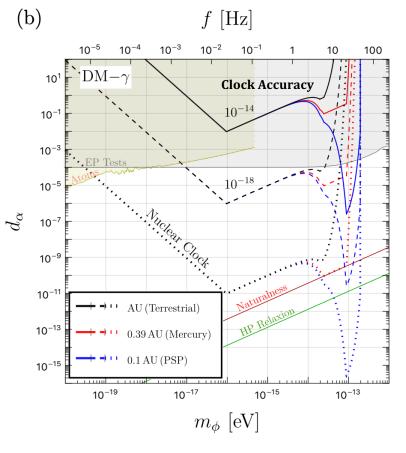
Enhancement of the DM Density



PSP: Parker Solar Probe Tsai, Eby, Safronova, arXiv:2112.07674

Results





- Motivate Specific Frequency Region!
- Motivate Nuclear Clocks!
- Tsai, Eby, Safronova, arXv:2112.07674

$$\mathcal{L} \supset \kappa \phi \left(d_{m_e} m_e \bar{e}e + \frac{d_\alpha}{4} F_{\mu\nu} F^{\mu\nu} + \frac{d_g \beta_3}{2g_s} G^A_{\mu\nu} G^{A\mu\nu} \right), \tag{1}$$

$$rac{g_e^2\Lambda^2}{(4\pi)^2} \lesssim m_\phi^2, \ \Lambda = 4\pi v_{EW} \simeq 3 \ {
m TeV}.$$

Naturalness condition

Relaxion Prediction

- For the Higgs portal-like theories, scalar couplings to matter are generated by mixing with the Higgs, and so can be parameterized by a relaxion φ-higgs mixing angle sinθ;
- One has $g_e = y_e \sin\theta$ and $g_{\gamma} \sim (\alpha/4\pi v) \sin\theta$, where y_e is the Higgs Yukawa coupling to the electron, v is the Electroweak vacuum expectation value.
- The green line is assuming maximum relaxion-higgs mixing, which is of order $g_e \sim y_e (m_{\phi}/m_H)$
- see, e.g., Banerjee, Budker, Eby, Kim, Perez, 1902.08212 for more discussions.

Spatial Variation of Fundamental Constants

$$k_X \equiv c^2 \frac{\delta X}{X \, \delta U}$$
. $X = \alpha, \mu, \text{ or } m_q / \Lambda_{QCD}$.

 δU : change in gravitational potential .

$$\delta U/c^2\simeq 3.3 imes 10^{-10},~$$
 Earth variation.

 $\delta U/c^2 \sim 9 imes 10^{-8},~$ from Earth to Solar probe at 0.1 AU.

• Achieve constraints on k_X that are a factor of ~ 300 stronger!

Space Mission & Telescopes



An artist's impression of the Lucy spacecraft performing a flyby of a Jupiter trojan.

NASA/SwRI and SSL/Peter Rubin https://www.nasa.gov/press-release/nasaselects-two-missions-to-explore-the-earlysolar-system

Lucy is a planned NASA space probe that will complete a 12-year journey to seven different **asteroids. Human landing?**



A photograph and rendering mix of the exterior of the Vera C. Rubin Observatory building on Cerro Pachón in Chile. Image credit: Rubin Obs./NSF/AURA

 Optical – Vera Rubin Observatory: increase the number of solarsystem objects by 5 times.



Gaia mapping the stars of the Milky Way

Optical – GAIA provides stellar reference for asteroid localization

Exciting Research Directions for Discussions

- Asteroidal/Planetary Tracking Array develop a tracking array to study bosonic ultralight dark matter (possible) and gravitational wave (difficult);
- 2. Probing dark energy and cosmology-motivated modified-gravity theories
- 3 Consider non-gravitational dark matter-SM interactions
- Quantum technologies in Space: <u>Q-SEnSE</u> + SpaceQ meeting
- See many relevant references in the backup slides



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Yu-Dai Tsai, UC Irvine, yudait1@uci.edu & yt444@cornell.edu

We can also discus & collaborate on ...

- Fixed-target searches for dark matter & long-lived particles (FerMINI & LongQuest) with Pospelov et al.
- LHC Forward Experiments: Forward Physics Facility, **FORMOSA** (a millicharge experiment I proposed), with Feng et al.
- Dark matter model building (dark sector QCD, Strongly Self-Interacting Dark Matter, SIMP/ELDER), with Murayama, Slatyer, Perelstein et al.
- Dark matter searches using neutron star / compact merger / multimessenger astronomy, with Profumo, Sathyaprakash et al.
- Neutrino physics (cosmic neutrino background) & neutrino BSM, with Shoemaker et al.
- Collaborating with **many awesome early-career collaborators.**

Invisible disabilities cultivate diverse abilities

Big Picture & Outlook

- Bridging planetary science, space (quantum) technology, and fundamental physics
- Our result is exciting now and has significant potential given the future measurements: radar, optical, and space missions will bring tremendous progress!
- Atomic clocks on the moon, spacecraft, satellite, Asteroid Tracking Array, and Advanced Lunar Ranging: Many exciting projects forward! Collaborating with NIST, NASA, ESA, etc people on proposals

Yu-Dai Tsai, UC Irvine, '21 <u>yt444@cornell.edu</u>

Let's protect the Earth & find dark matter; happy to discuss more

Thank you!

Thank Josh, Marianna, Jason, Luca, Sunny, Youjia for comments My outreach interview: <u>https://www.youtube.com/watch?v=xDX9XwLHBuM</u> (> 76K views!)

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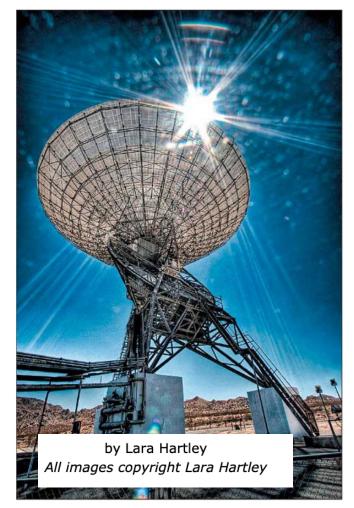
More References

- Seto, Cooray, arXiv:0405216, PRD 04
- LLR Experiments: Williams, Turyshev, Boggs, PRL 04 Murphy, Rept. Prog. Phys 13
- Atomic / nuclear clocks for fundamental physics: Peik, Schumm, Safronova, Pálffy, Weitenberg, Thirolf, 2012.09304
- GW background, Fedderke, Graham, Rajendran, PRD21
 GW measurement with atomic clocks, Fedderke, Graham, Rajendran, 2112.11431
- Quantum Technologies in Space, Kaltenbaek, Exp Astron 21

Radar Observations

- Radar Goldstone Observatory: Provide very precise location and velocity information of the asteroids
- Radar astronomy: observing nearby astronomical objects by reflecting microwaves off target objects and analyzing the reflections.
- Round-trip light time (RTLT): The elapsed time taken by a signal travelling from the Earth to a spacecraft or other celestial body
- Doppler shift:





Students can control the huge Echo radio telescope to collect data from objects in the universe at which the antenna is pointed.

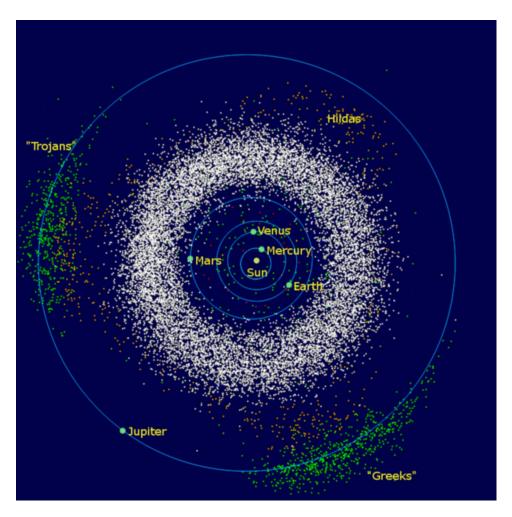
https://www.desertusa.com/desertcalifornia/goldstone-deep-space.html

Asteroids



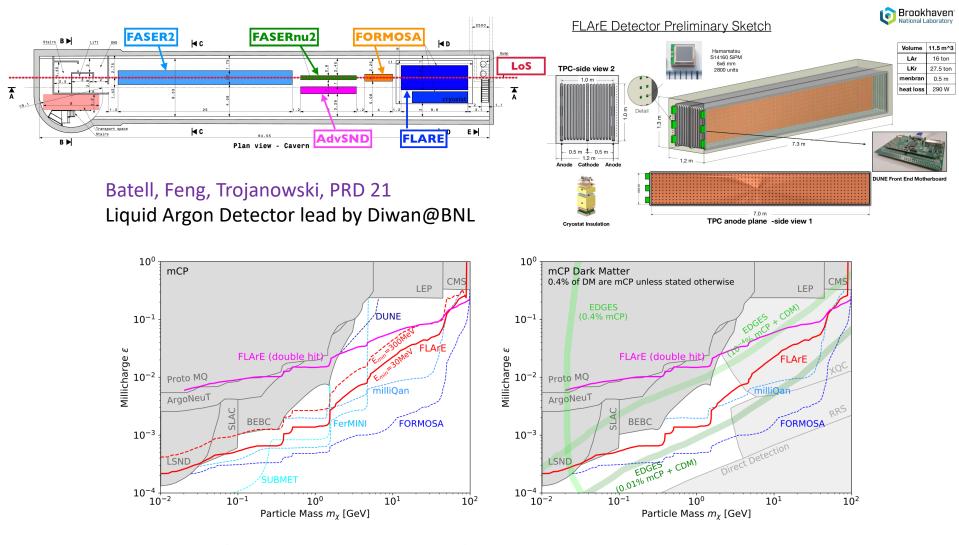
"Is he not the celebrated author of *The Dynamics of an Asteroid*, a book which ascends to such rarefied heights of pure mathematics that it is said that there was no man in the scientific press capable of criticizing it? — *Sherlock Holmes, The Valley of Fear*

"The mor hazardous the asteroids, the better for fundamental Physics" -- Professor Moriarty (maybe)



https://commons.wikimedia.org/wiki/File:InnerSolarSystemen.png, public domain, granted usage for any purposes

FLArE up millicharges & EM-form factor dark sector



Kling, Kuo, Trojanowski, Tsai, arXiv:2205.09137